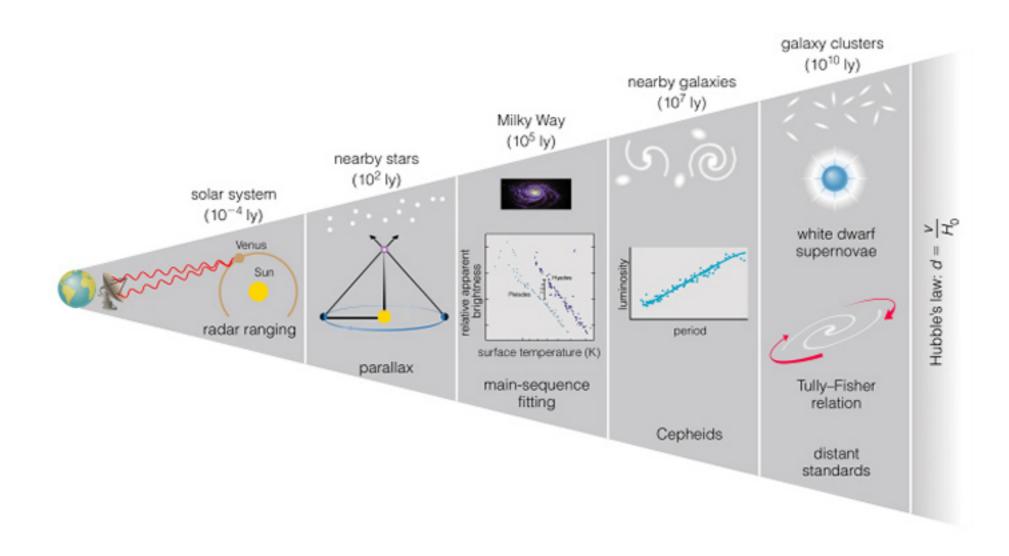
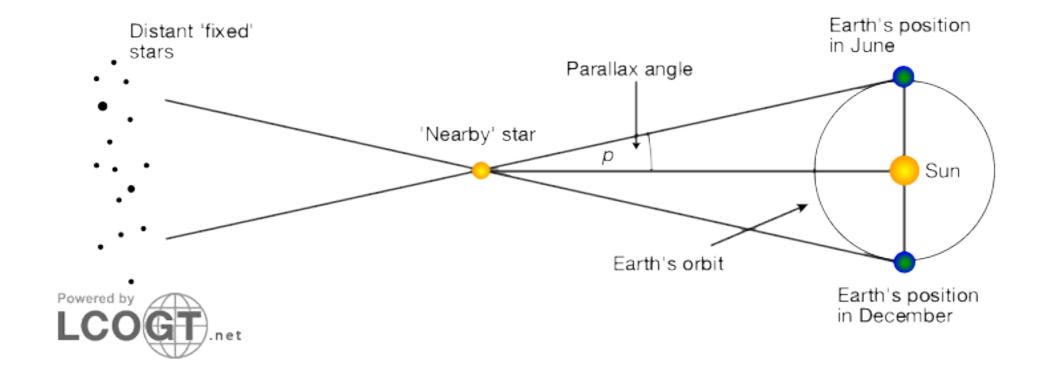
Cosmic Distance Ladder

= succession of methods by which astronomers determine the distance to celestial objects



Parallax



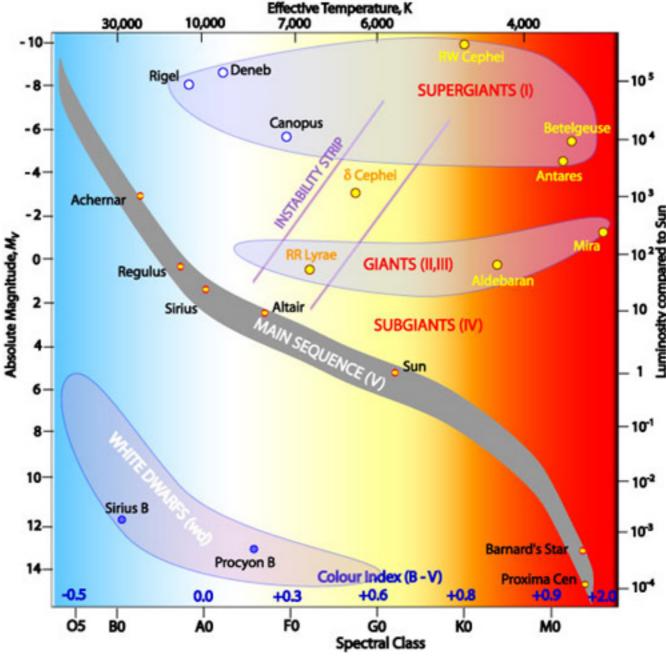
Def. Parsec:
$$\frac{1 \text{ Au}}{1 \text{ pc}} = 1'' = 1 \text{ arcsec}$$
 $\frac{1 \text{ pc} = 3.26 \text{ lyr}}{= 3.09 \cdot 10^{16} \text{ m}}$

With the Hubble telescope: Can reach up to 5000 parsec.

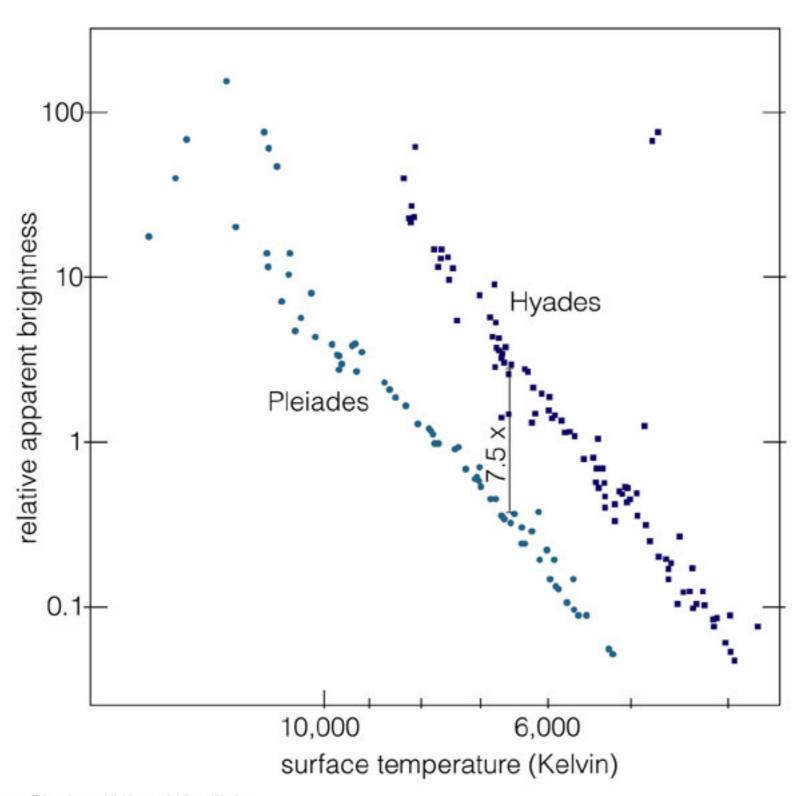
Hertzsprung-Russell diagram

"Main sequence stars" have a known relation between temperature (spectral class) and luminosity!

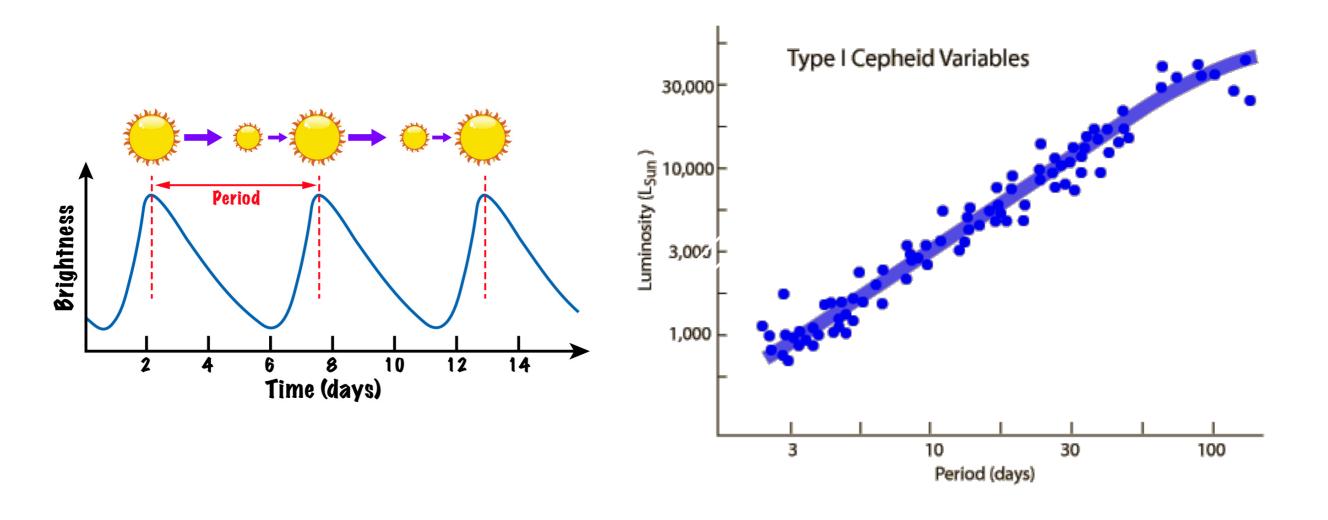
Hertzsprung-Russell Diagram



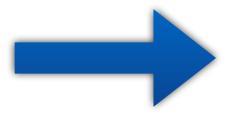
Main sequence fitting



Cepheids = periodically pulsating stars

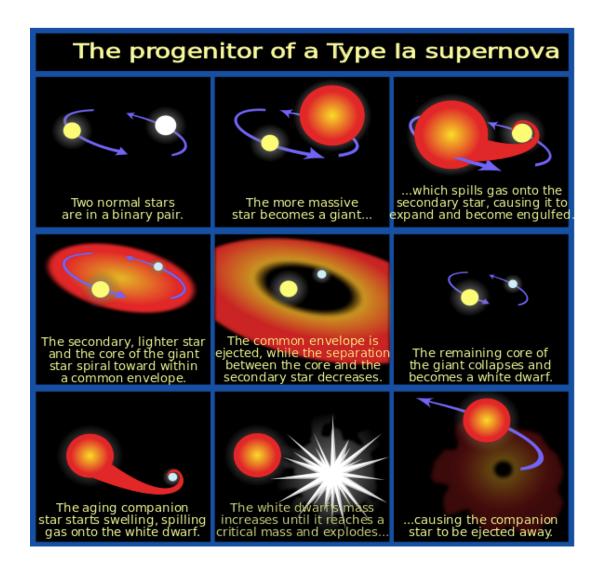


Known luminosity (when period is measured)

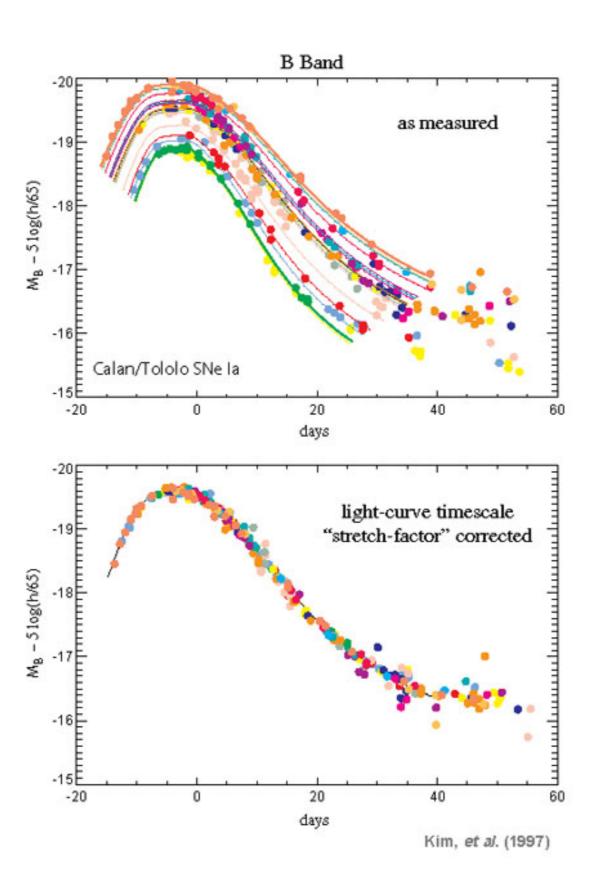


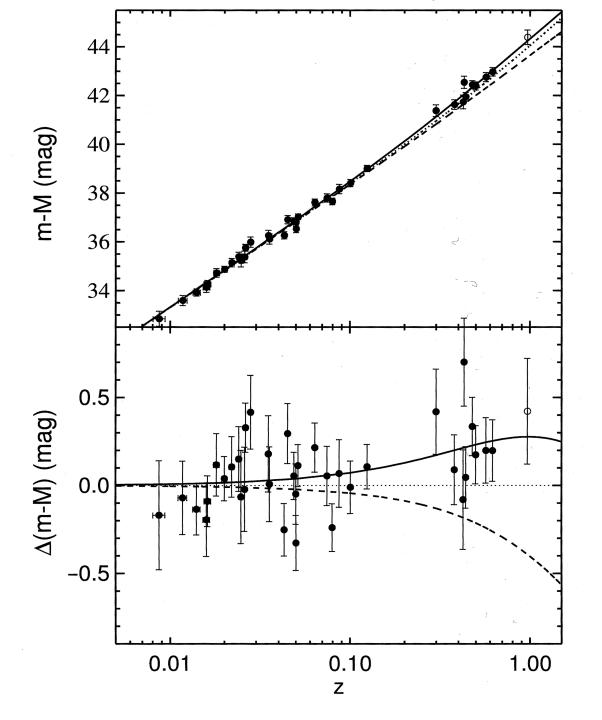
Standard candle

The ultimate standard candles for large luminosity distances: Supernovae type la



Supernovae of Type la





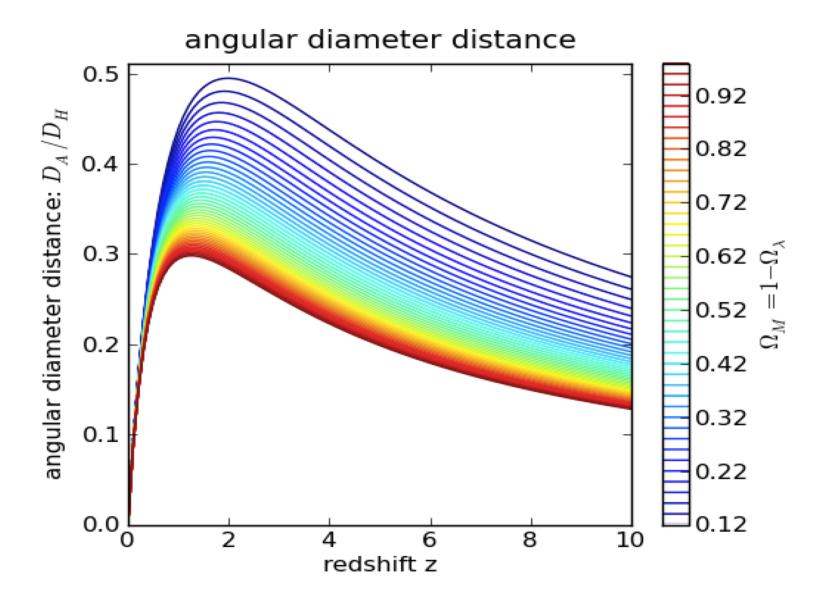
Top: Hubble diagram with Supernovae type la:

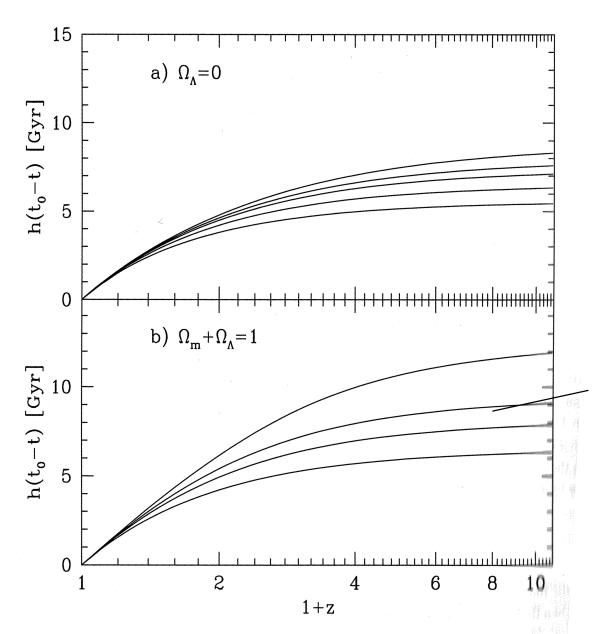
Distance modulus m – M = 5log(d_L/10pc) -uses "astronomer" units rather than "physicist" units (eg Mpc)

Points= observations Curves= fits with different cosmological models (best fit = LCDM model)

Bottom: residuals of data points from fits

Solid line shows Accepted model (LCDM): Omega_ = 0.27, Omega_Lambda = 0.73





Lookback time vs. Redshift for different cosmological models

In both plots from top to bottom we have increasing contribution of Omega_m

Accepted model (LCDM): Omega_ ~ 0.3, Omega_Lambda ~ 0.7

Note age of the Universe (=asymptotic limit of curves) increases for increasing Omega_Lambda

Depends also on "h", which Is measured to be ~ 0.7